

Wake of a Charge Moving in the Vicinity of a Dielectric Structure

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Outline

- Motivation
- Configurations
- Charged Line & Dielectric Cylinder
- Formulation
- Simulation Results
- Summary

Motivation

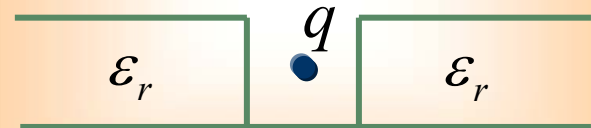
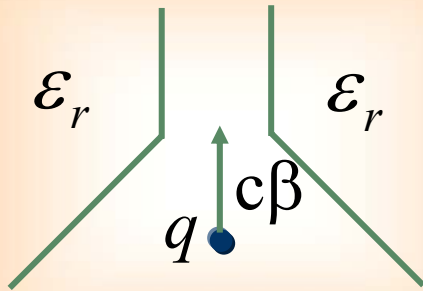
- ◆ Vacuum acceleration with lasers entails *relativistic* motion of a bunch in the vicinity of metallic or dielectric structures
- ◆ *Non-relativistic* electrons are used for ultra-microscopy of surfaces at nano-meter scale
- ◆ It is necessary to determine the characteristics of the *wakes* generated by a bunch in a variety of configurations

Motivation: Applications

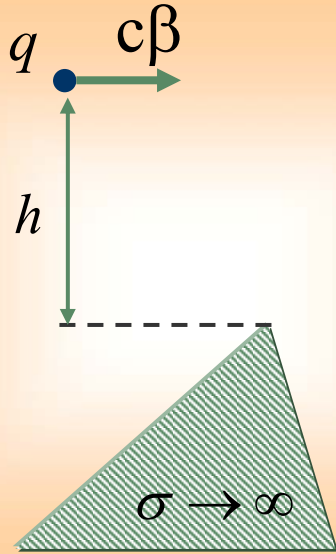
- ◆ Ultra-microscopy (Chemistry, Biology)
- ◆ X-ray sources (Medicine, Biology, Chemistry, Physics)
- ◆ Spectroscopy (Biology, Chemistry)
- ◆ Particle accelerators (Medicine, Physics)

Configurations of Interest

Point Charge & Accelerating Structure

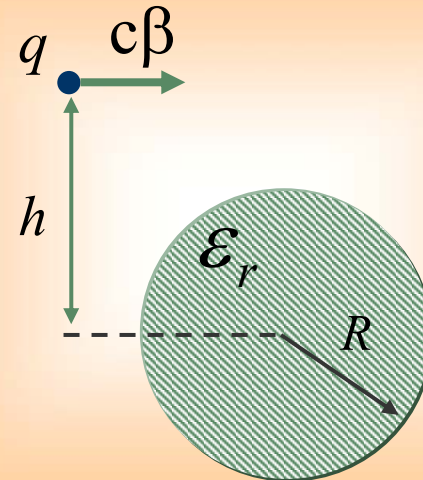


Configurations of Interest

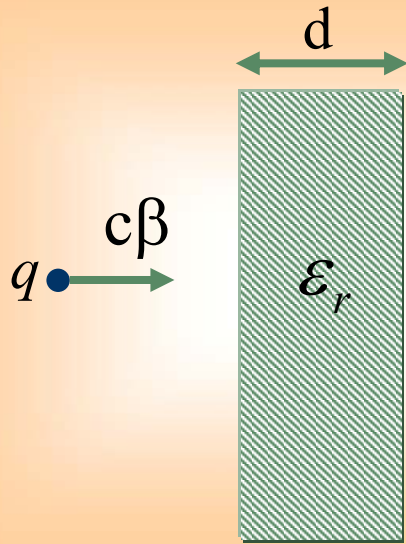


Point Charge & Metallic Edge

Point Charge & Dielectric Cylinder

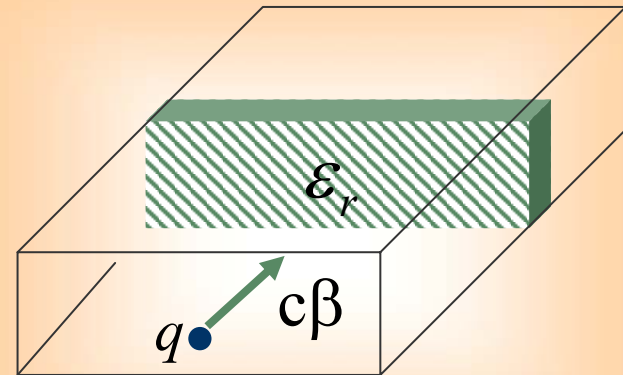


Configurations of Interest



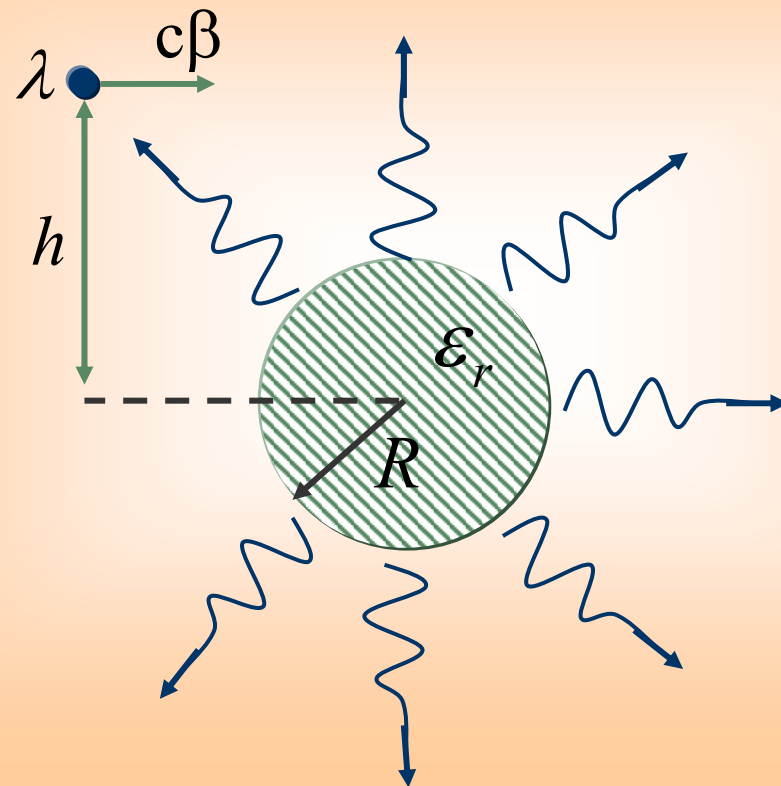
Point Charge & Dielectric Layer

Point Charge & Loaded Wave-Guide



Configuration:

Charged-Line & Dielectric Cylinder



$$\frac{\partial}{\partial z} = 0$$

Formulation: Primary Field

$$J_x(x, y; \omega) = -\frac{\lambda}{2\pi} \delta(y - h) \exp(-j \frac{\omega}{c\beta} x)$$

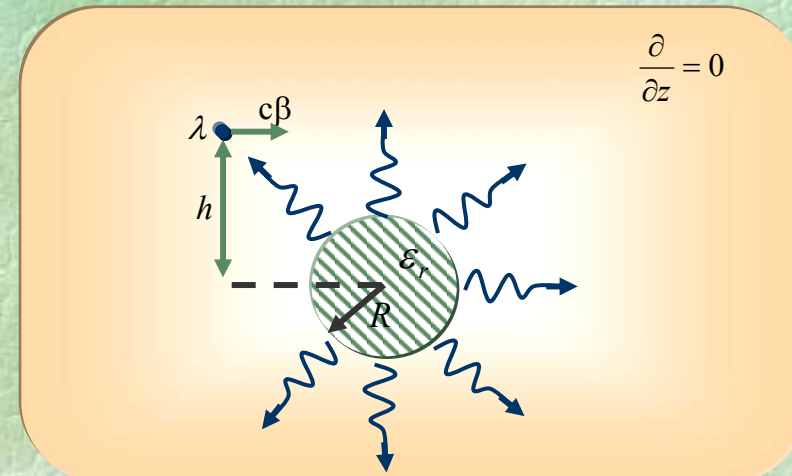
$$E_\varphi^{(p)} = \frac{-\lambda \mu_0}{4\pi \Gamma} \exp\left(-j \frac{\omega}{c\beta} r \cos \varphi - \frac{\omega}{c\gamma\beta} |r \sin \varphi - h|\right) \left[\frac{-j\omega}{\gamma^2 \beta^2} \sin \varphi + \frac{\omega}{\gamma \beta^2} \operatorname{sgn}(y - h) \cos \varphi \right]$$

$$H_z^{(p)} = \frac{-\lambda}{4\pi} \exp\left(-j \frac{\omega}{c\beta} r \cos \varphi - \frac{\omega}{c\gamma\beta} |r \sin \varphi - h|\right) \operatorname{sgn}(y - h)$$

$$\gamma = \frac{1}{\sqrt{1 - \beta^2}}; \quad \Gamma = \frac{\omega}{c} \frac{1}{\gamma \beta}$$

$$x = r \cos \varphi$$

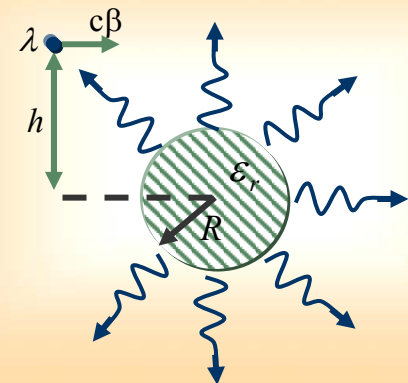
$$y = r \sin \varphi$$



Formulation: Secondary Field

$$H_Z^{(s)}(r, \varphi; \omega) = \sum_{n=-\infty}^{\infty} e^{jn\varphi} \begin{cases} A_n H_n^{(2)}\left(\frac{\omega}{c} r\right) & r > R \\ B_n J_n\left(\frac{\omega}{c} \sqrt{\epsilon_r} r\right) & r < R \end{cases}$$

$$E_\varphi^{(s)}(r, \varphi; \omega) = -\frac{1}{j\omega\epsilon_0} \sum_{n=-\infty}^{\infty} e^{jn\varphi} \begin{cases} A_n \frac{\omega}{c} \dot{H}_n^{(2)}\left(\frac{\omega}{c} r\right) & r > R \\ \frac{B_n}{\sqrt{\epsilon_r}} \frac{\omega}{c} j_n\left(\frac{\omega}{c} \sqrt{\epsilon_r} r\right) & r < R \end{cases}$$



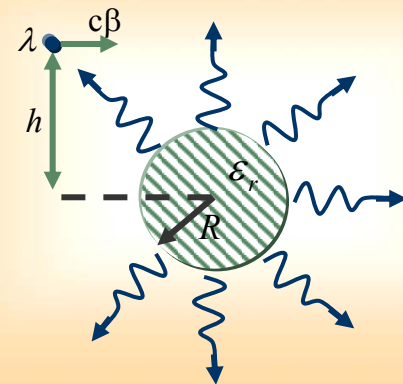
$$\frac{\partial}{\partial z} = 0$$

Formulation: Emitted Energy

$$P(r,t) = 2\pi r \frac{1}{2\pi} \int_{-\pi}^{\pi} d\varphi E_{\varphi}^{(s)}(r, \varphi; t) H_z^{(s)}(r, \varphi; t)$$

$$W(r) = \int_{-\infty}^{\infty} dt P(r, t)$$

$$r \gg \lambda \quad \Rightarrow \quad W = \frac{16\pi}{\epsilon_0} \int_0^{\infty} d\omega \frac{1}{\omega} \left[\sum_{n=-\infty}^{\infty} |A_n(\omega)|^2 \right]$$



$$\frac{\partial}{\partial z} = 0$$

Formulation: Field Solution

$$\Omega = \frac{\omega}{c} h$$

$$\bar{A}_n = \frac{A_n}{\lambda / 4\pi}$$

$$\bar{W} = \frac{W}{\lambda^2 / \pi \epsilon_0}$$

$$\bar{W} = \int_0^\infty d\Omega \frac{1}{\Omega} \sum_{n=-\infty}^\infty |\bar{A}_n(\Omega)|^2 \quad \frac{d\bar{W}}{d\Omega} = \frac{1}{\Omega} \sum_{n=-\infty}^\infty |\bar{A}_n(\Omega)|^2$$

$$\bar{A}_n(\Omega) = \frac{\left[\frac{V_n}{j\beta} \sqrt{\epsilon_r} J_n \left(\sqrt{\epsilon_r} \Omega \frac{R}{h} \right) - U_n j_n \left(\sqrt{\epsilon_r} \Omega \frac{R}{h} \right) \right] e^{-\frac{\Omega}{\gamma\beta}}}{H_n^{(2)} \left(\Omega \frac{R}{h} \right) j_n \left(\sqrt{\epsilon_r} \Omega \frac{R}{h} \right) - \sqrt{\epsilon_r} \dot{H}_n^{(2)} \left(\Omega \frac{R}{h} \right) J_n \left(\sqrt{\epsilon_r} \Omega \frac{R}{h} \right)}$$

$$U_n = \left(\frac{\gamma+1}{\gamma-1} \right)^{n/2} J_n \left(\Omega \frac{R}{h} \right) e^{-jn\pi/2}; \quad V_n = \frac{1}{2} \left(1 + \frac{j}{\gamma} \right) U_{n-1} + \frac{1}{2} \left(1 - \frac{j}{\gamma} \right) U_{n+1}$$

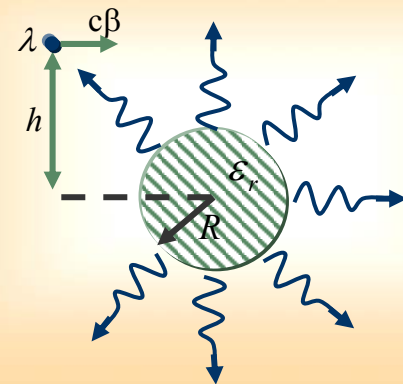
Formulation: Longitudinal Impedance

$$Z_{\parallel} = -\frac{1}{\lambda} \int_{-\infty}^{\infty} dx E_x(x, h; \omega) \exp(j \frac{\omega}{c\beta} x)$$

$$W = -\frac{\lambda}{2\pi} \int_{-\infty}^{\infty} d\omega \int_{-\infty}^{\infty} dx E_x(x, h; \omega) \exp(j \frac{\omega}{c\beta} x)$$

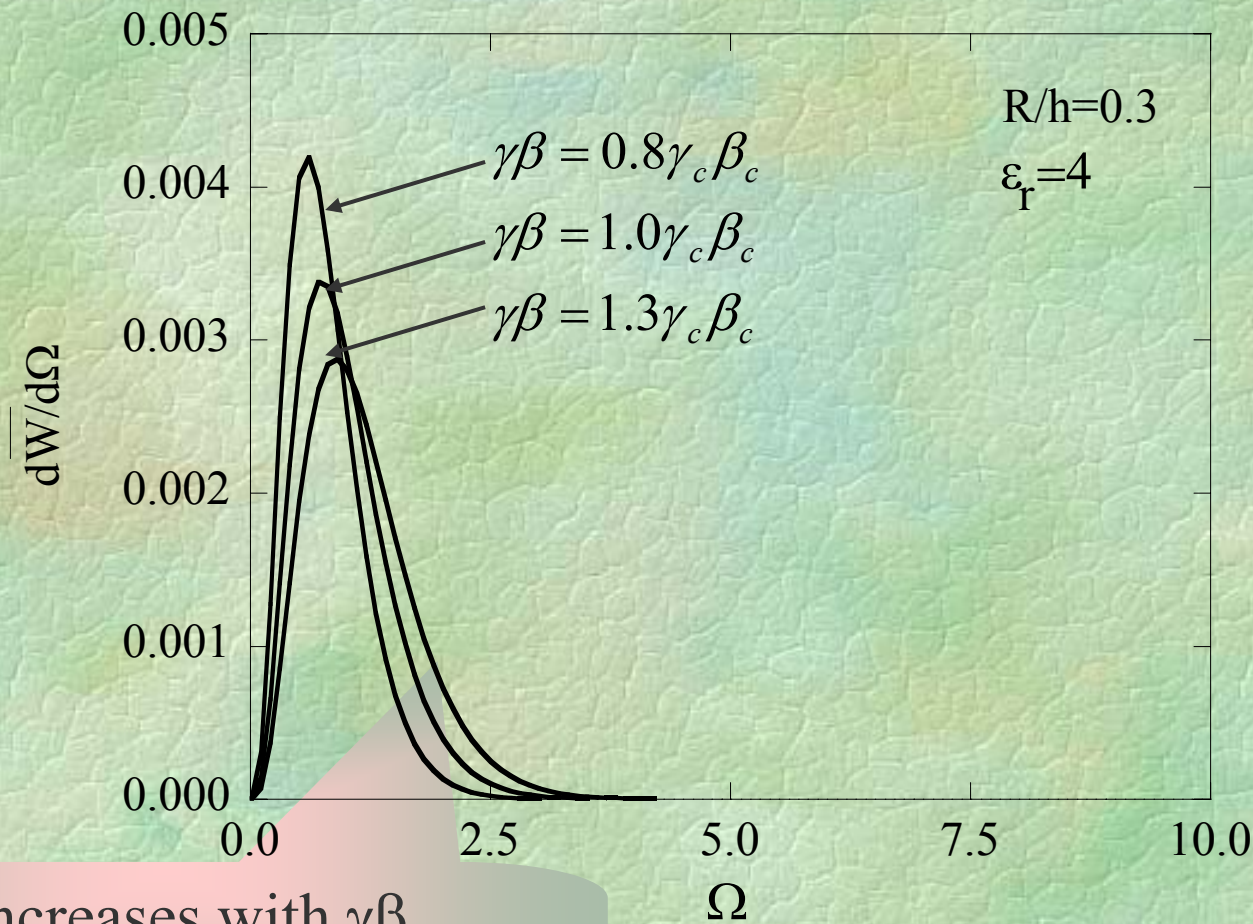
$$\Rightarrow \frac{dW}{d\omega} = \frac{\lambda^2}{2\pi} Z_{\parallel}$$

$$\bar{Z}_{\parallel} \equiv \frac{1}{2\eta_0 h} Z_{\parallel} \quad \Rightarrow \quad \frac{d\bar{W}}{d\Omega} = \bar{Z}_{\parallel}$$



$$\frac{\partial}{\partial z} = 0$$

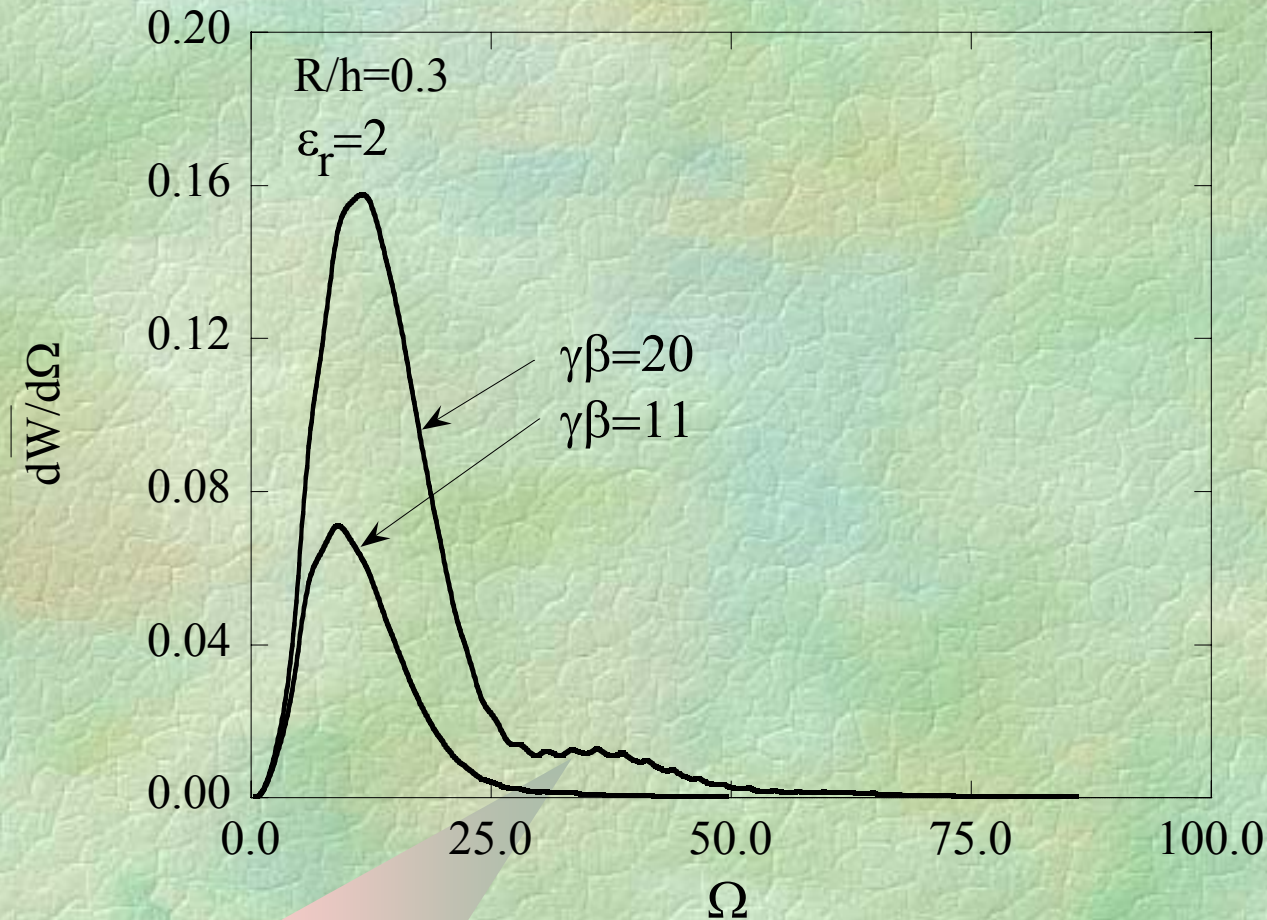
Longitudinal Impedance: Low Energy



Ω_{peak} increases with $\gamma\beta$

Narrow spectrum

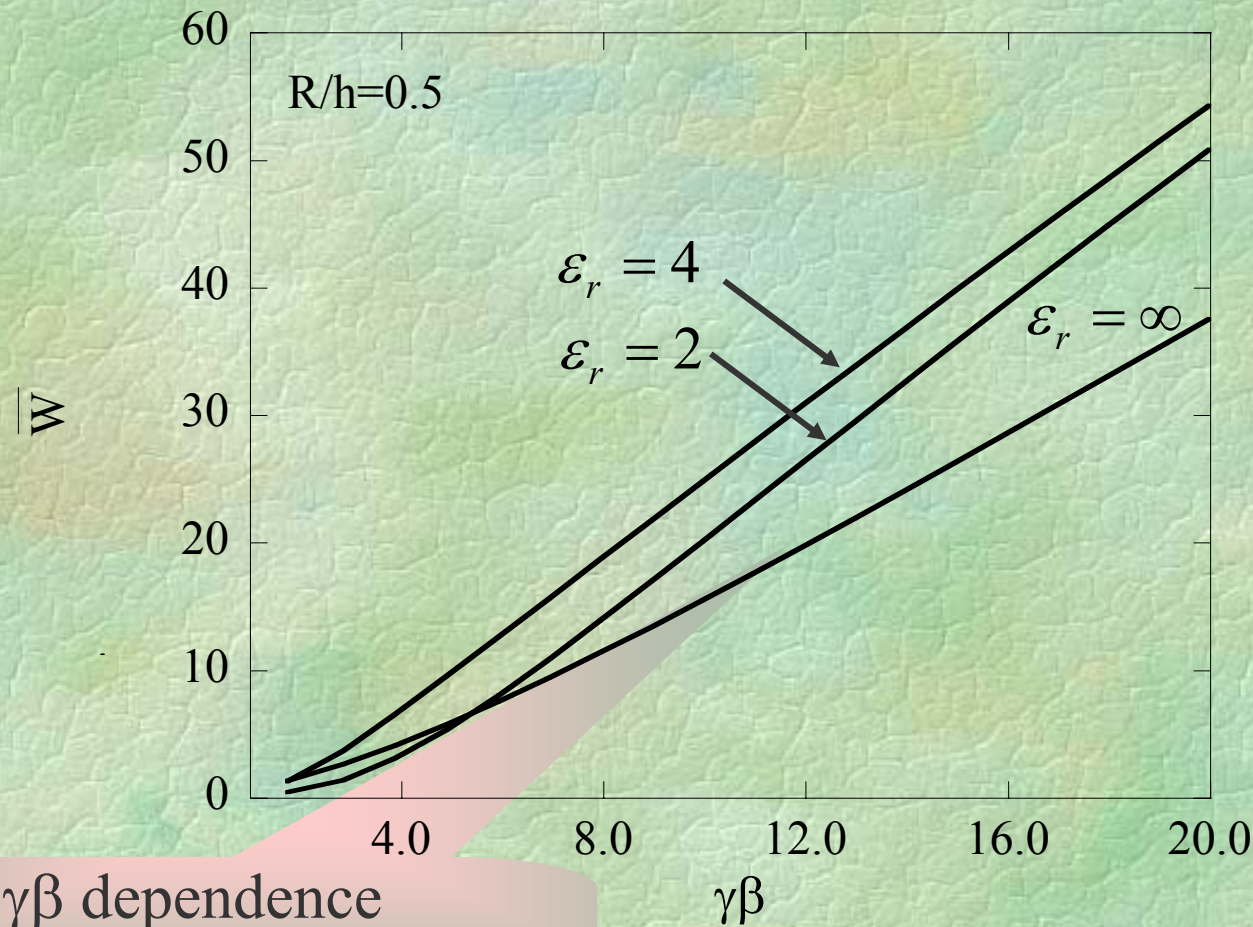
Longitudinal Impedance: High Energy



Ω of peak increases with $\gamma\beta$

Wide spectrum

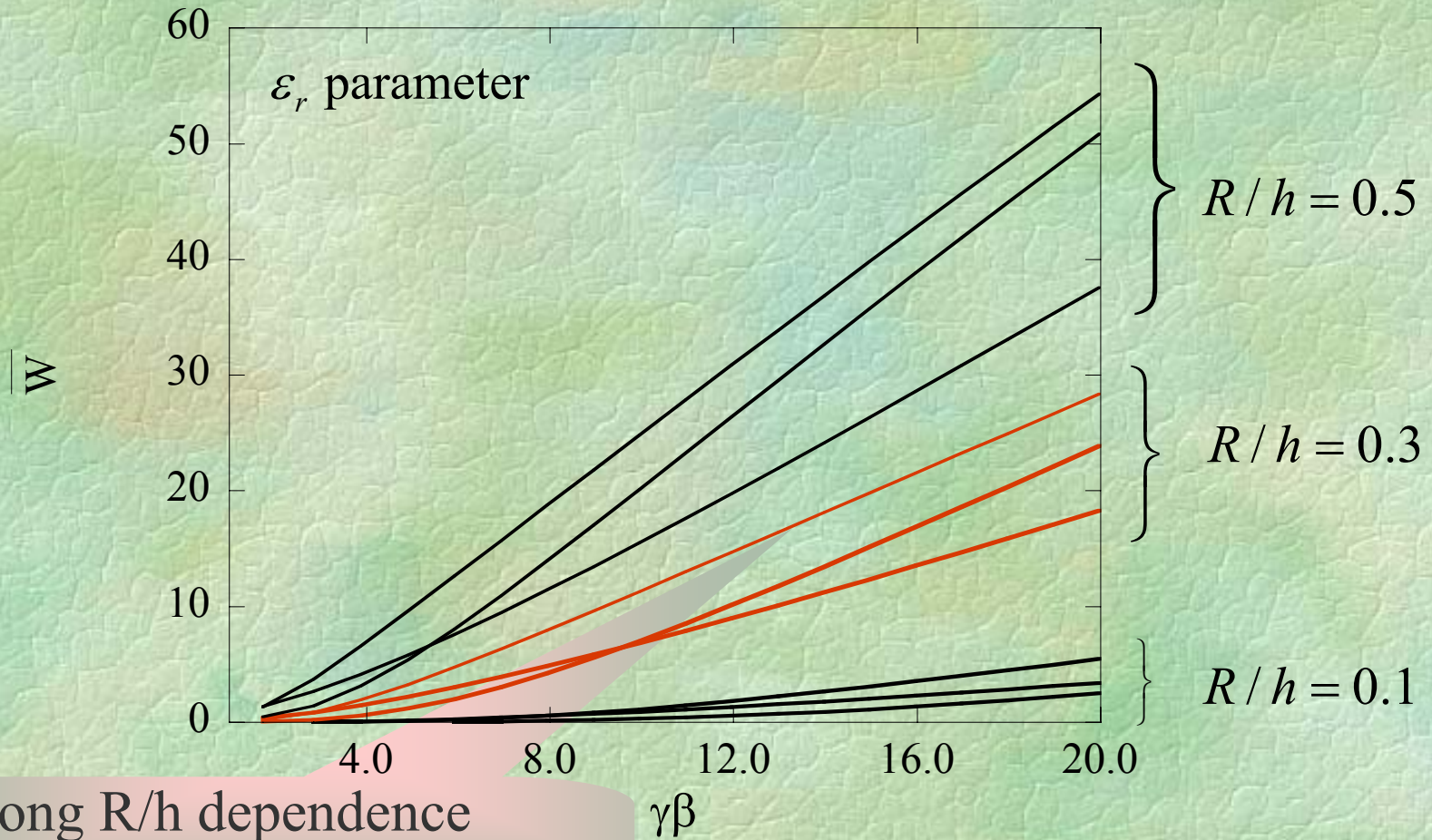
Radiated Energy: Relativistic Case



Linear $\gamma\beta$ dependence

$\overline{W}_\infty < \overline{W}$ for $\gamma\beta \gg 1$

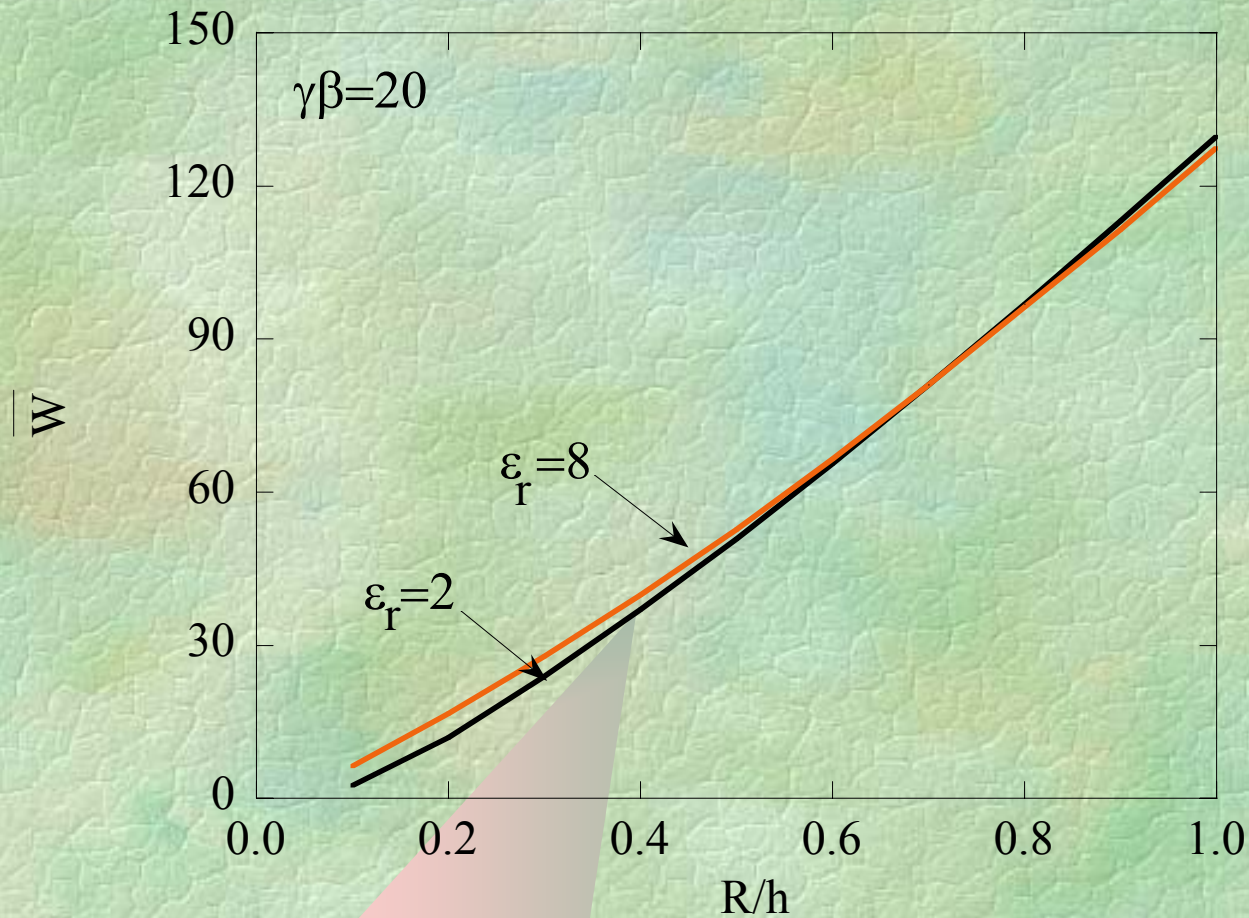
Radiated Energy: Relativistic Case



Strong R/h dependence

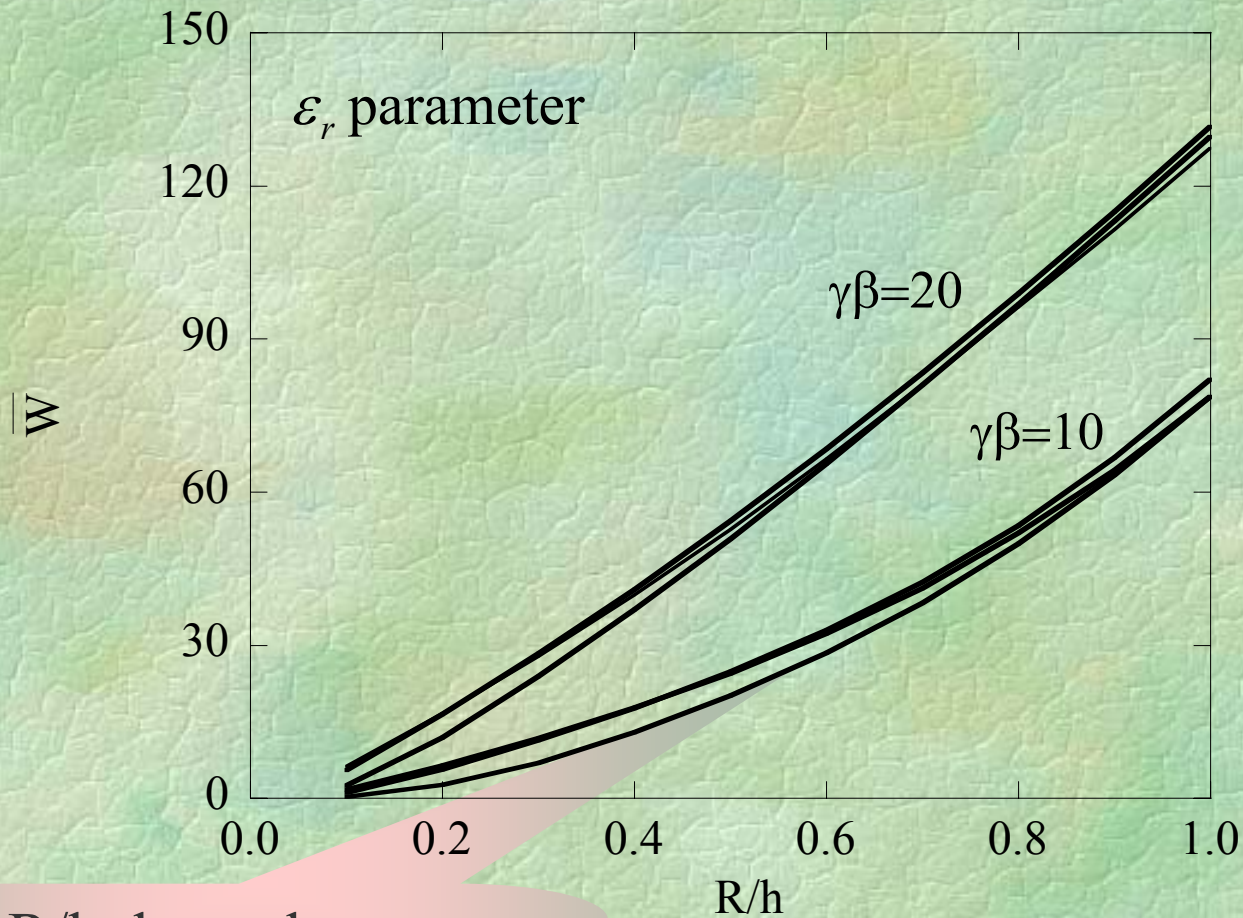
Linear $\gamma\beta$ dependence

Radiated Energy: Relativistic Case



Almost linear R/h dependence

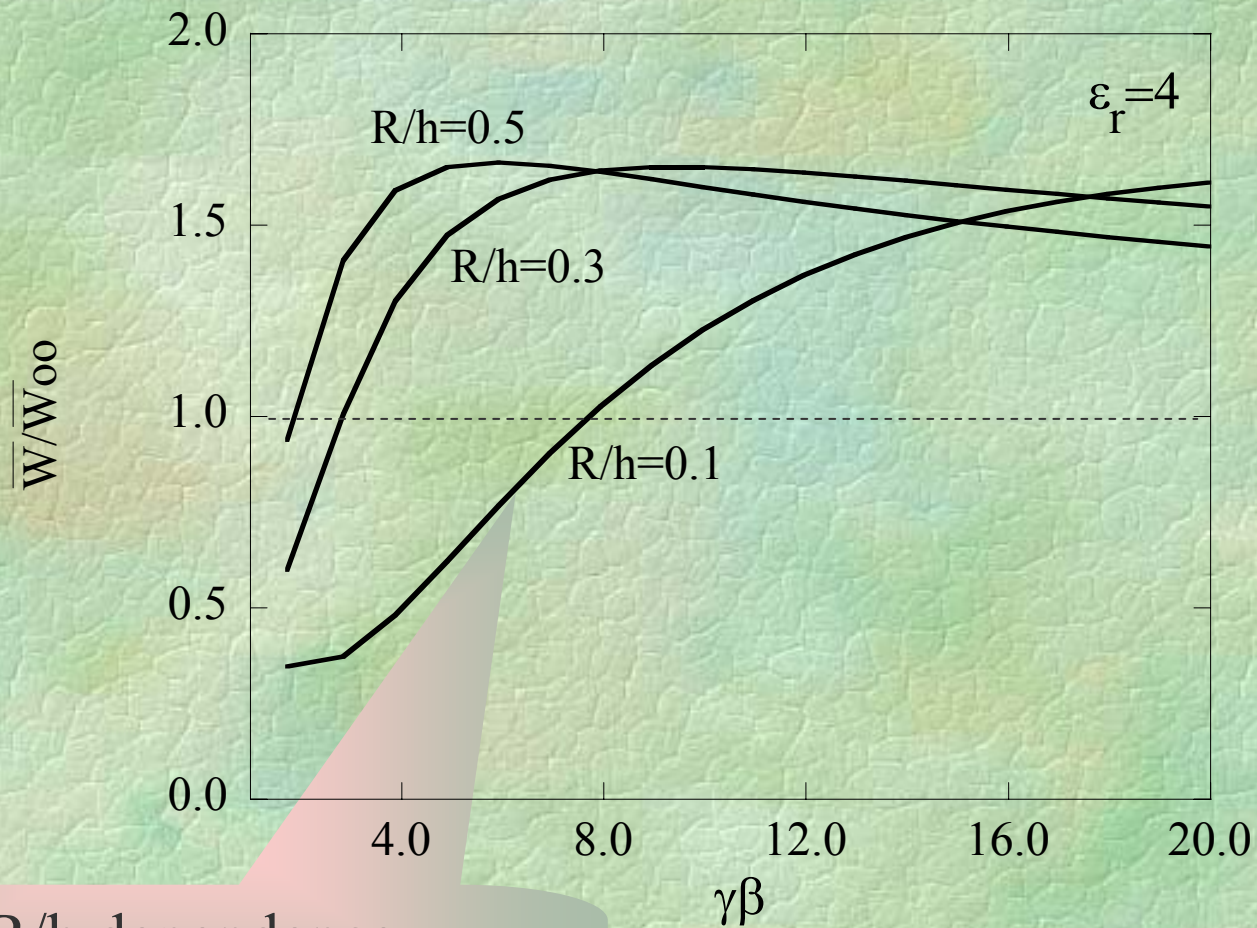
Radiated Energy: Relativistic Case



Strong R/h dependence

Weak ϵ_r dependence

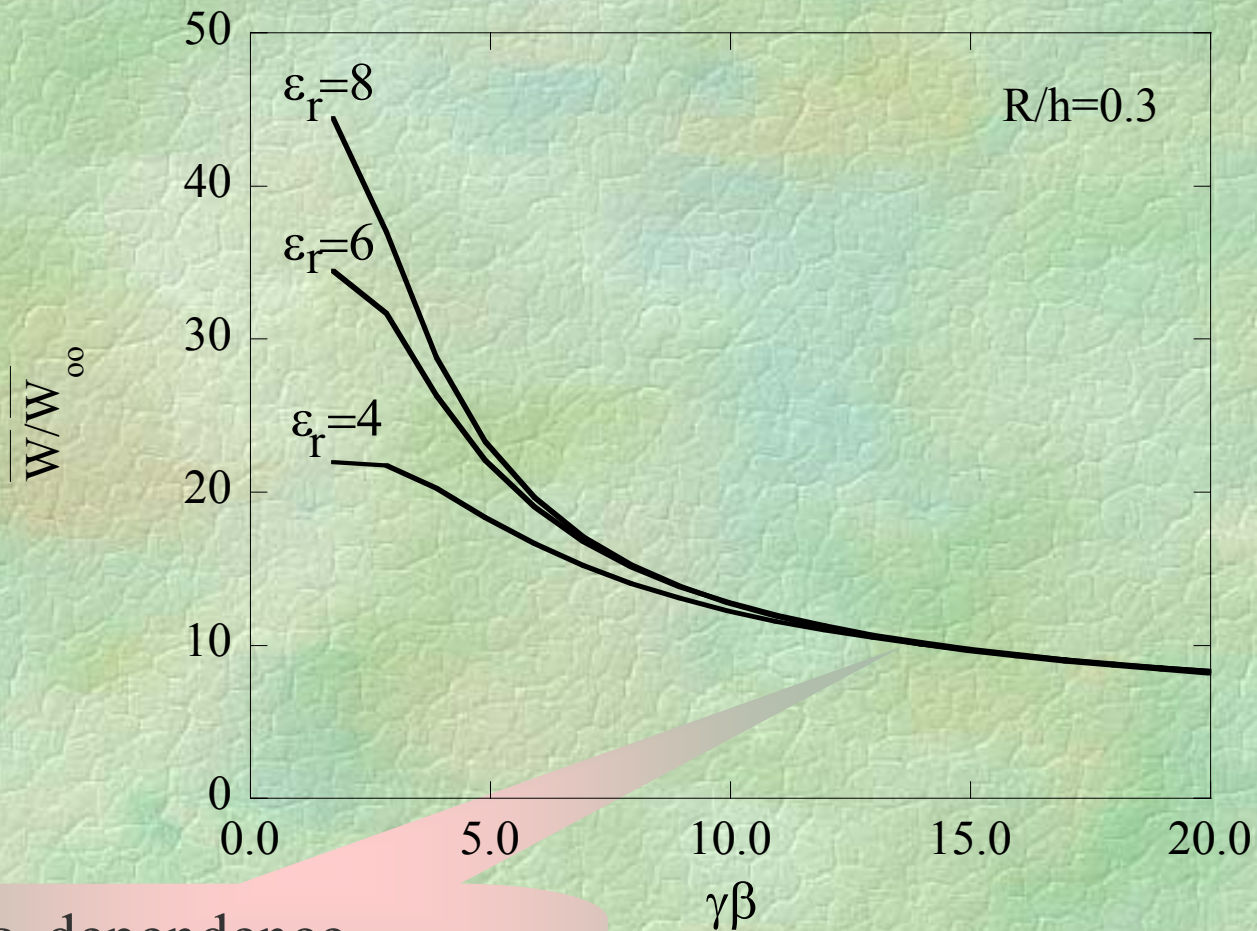
Radiated Energy: Relativistic Case



Weak R/h dependence

Weak $\gamma\beta$ dependence

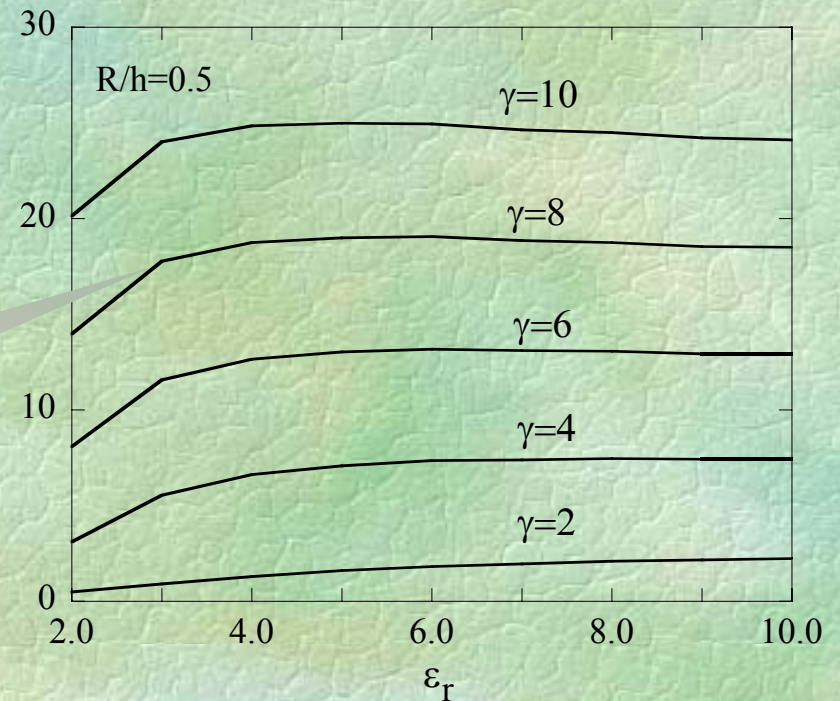
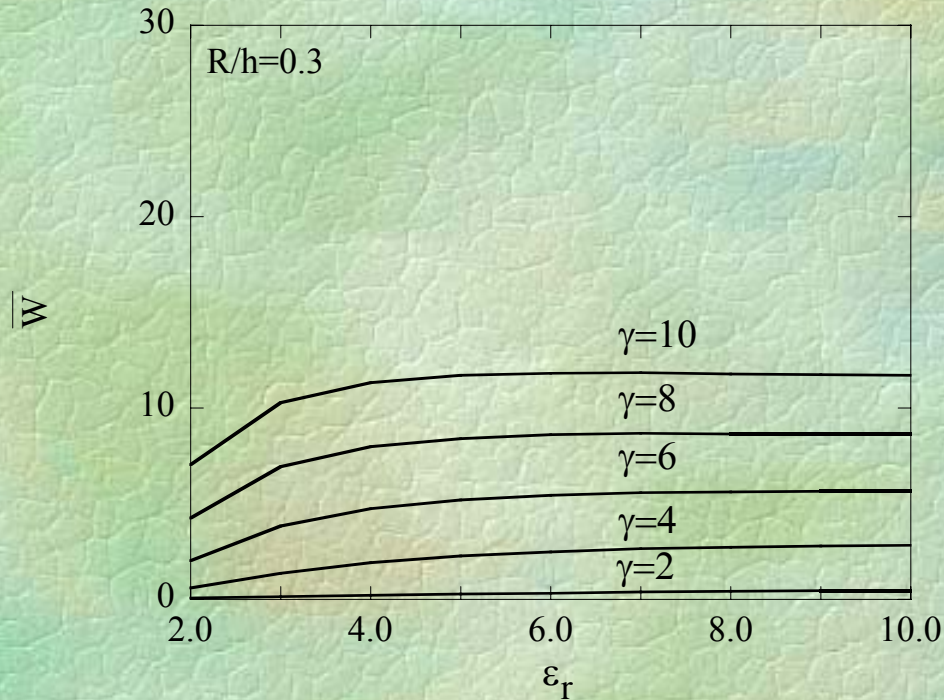
Radiated Energy: Relativistic Case



Weak ϵ_r dependence

Strong $\gamma\beta$ dependence

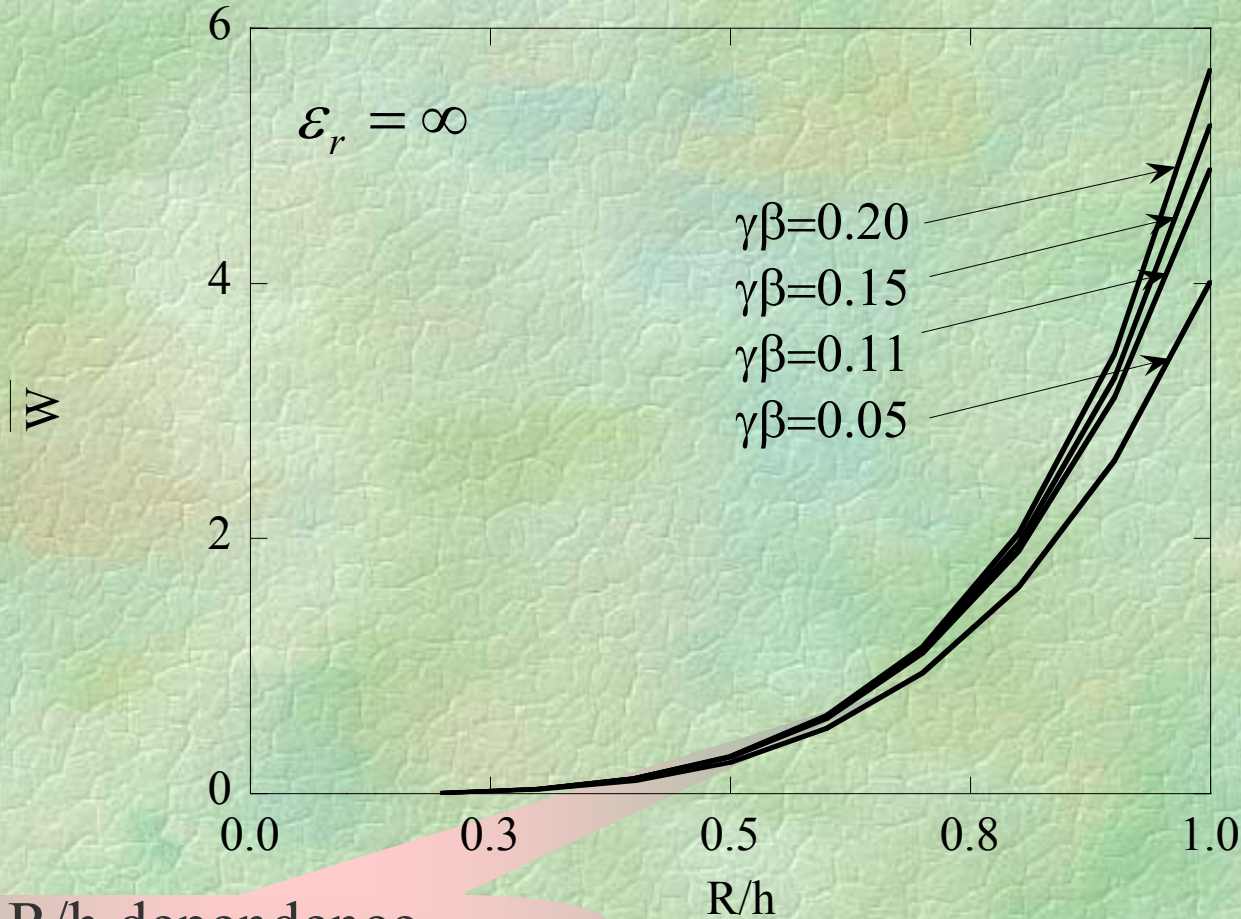
Radiated Energy: Relativistic Case



Weak ϵ_r dependence

Strong $\gamma\beta$ dependence

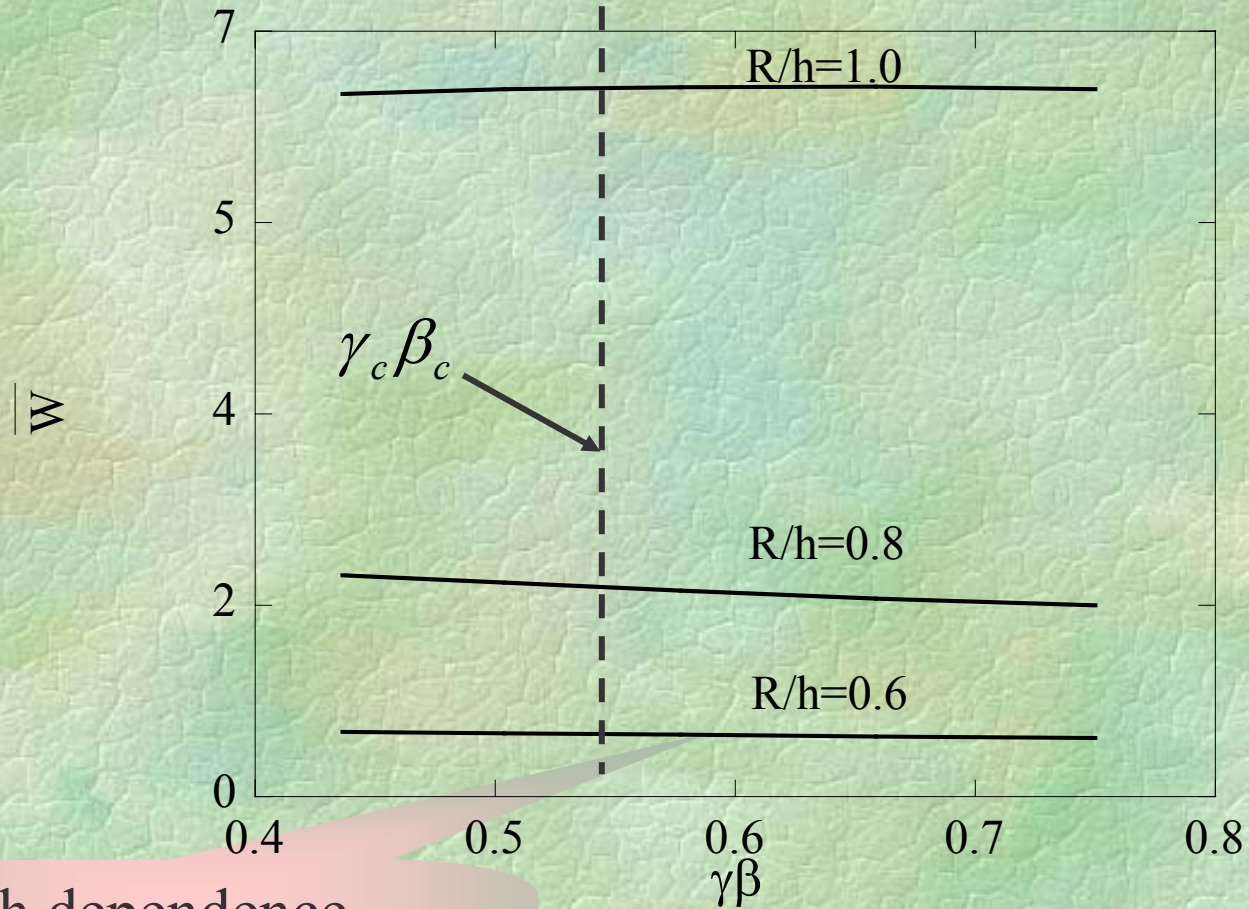
Radiated Energy: Low Energy



Strong R/h dependence

Weak $\gamma\beta$ dependence

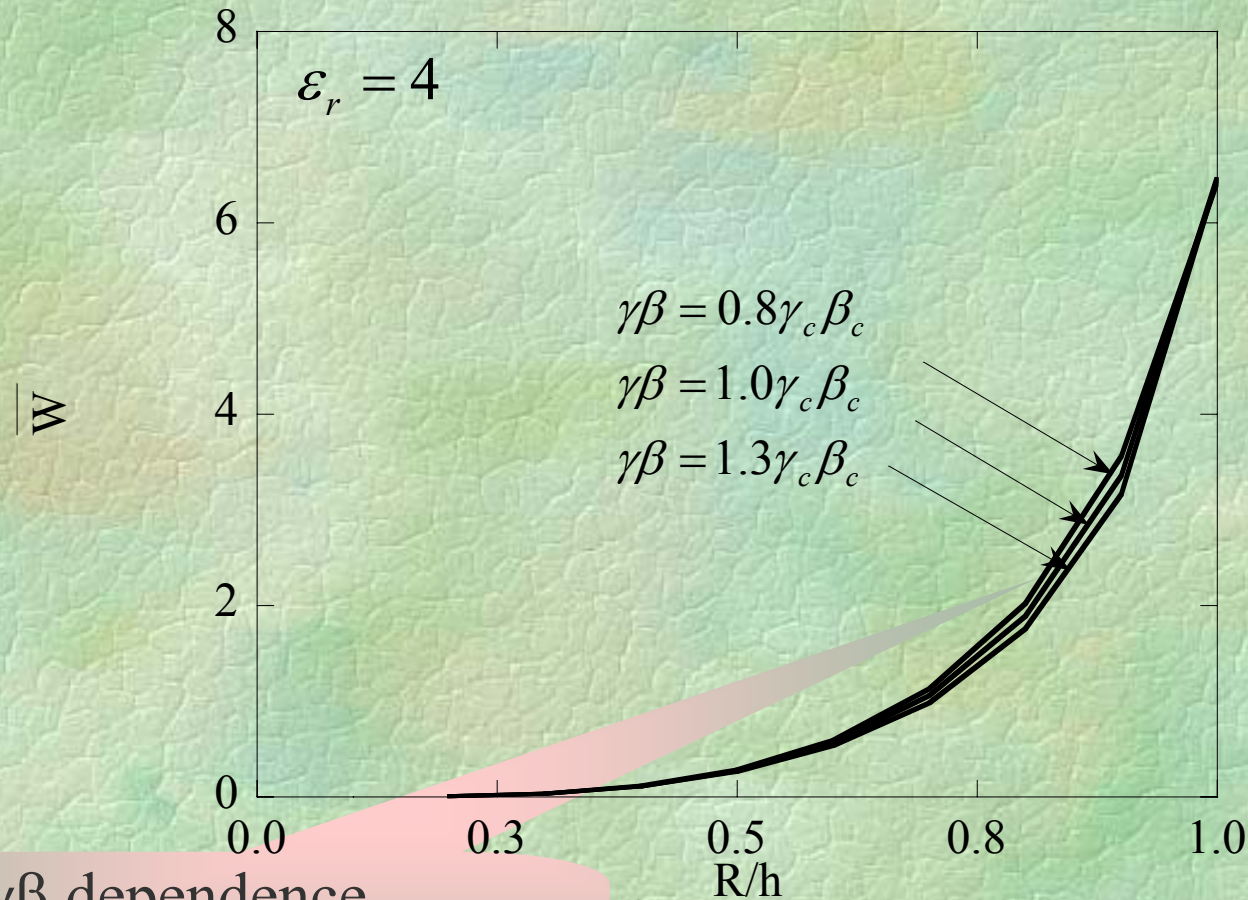
Radiated Energy: Transition & Cerenkov



Strong R/h dependence

Weak $\gamma\beta$ dependence

Radiated Energy: Transition & Cerenkov



Weak $\gamma\beta$ dependence

Strong R/h dependence

Summary

- Peak of the spectrum (Ω_{peak}) increases with $\gamma\beta$
- Spectrum width ($\Delta\Omega$) increases with $\gamma\beta$
- $\overline{W}$ is linear with $\gamma\beta$ for $\gamma\beta \gg 1$
- $\overline{W} > \overline{W}_\infty$ for $\gamma\beta \gg 1$
- $\overline{W}$ is linear with R/h for $\gamma\beta \gg 1$
- $\overline{W}$ virtually saturates as a function of ε_r
- $\overline{W}$ does not change significantly at the Cerenkov point
- $\overline{W} / \overline{W}_\infty$ has a maximum as a function of $\gamma\beta$